

Bayesian Neural Networks

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MLG reading group – 22nd Feb 2023



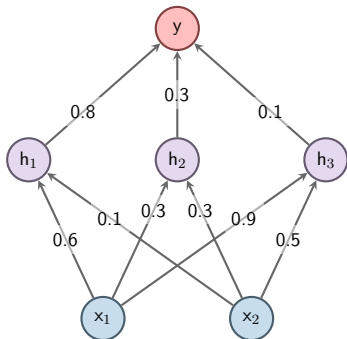
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Outline

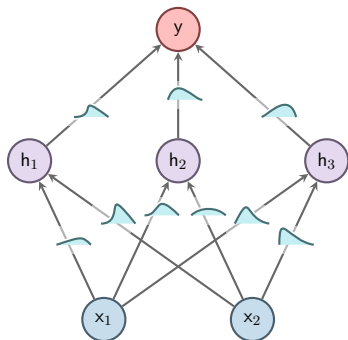
- Part 1 – James
 - BNN introduction
 - BNN challenges and solutions
 - BNN properties
- Part 2 – Javier
 - Laplace approximation
 - linear models
 - connections to infinite width limits
- Part 3 – Vincent
 - BNN priors
 - weight space
 - function space

Part 1

What are Bayesian Neural Networks?

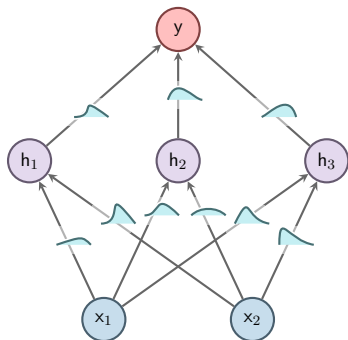


What are Bayesian Neural Networks?



$$p(\theta | \mathcal{D}) = \frac{p(\mathcal{D} | \theta) p(\theta)}{\int_{\theta} p(\mathcal{D} | \theta) p(\theta) d\theta} \quad (1)$$

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$$p(\mathbf{y}^* | \mathbf{x}^*) = \int_{\theta} p(\mathbf{y}^* | \mathbf{x}^*, \theta) p(\theta | \mathcal{D}) d\theta \quad (2)$$

Why BNNs?

- ① NNs are poorly calibrated – they don't know when they don't know!
 - “Reject” uncertain predictions.
 - Exploration in RL / Bandits.
 - Active learning.
 - Combining different model's predictions.
 - Bet sizing.
 - etc.
- ② Choosing hyper-parameters in NNs is hard (or expensive).
- ③ NNs can't naturally deal with missing data.

These are all problems that are solved by principled probabilistic models.

BNNs are hard

Some challenges:

- Integration!
- Choosing priors. (Part 3)
- High dimensionality.

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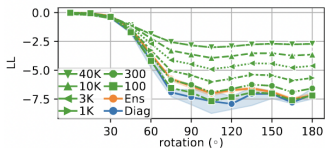
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- Choosing priors. (Part 3)
- High dimensionality.
 - Even *smaller* modern NNs have many parameters $> \mathcal{O}(10^6)$!
 - Storage of covariance matrices requires $\mathcal{O}(N^2)$ memory.
 - Makes approximation difficult.
 - Subspace [Izmailov et al., 2020] and subnetwork [Daxberger et al., 2021b] inference.

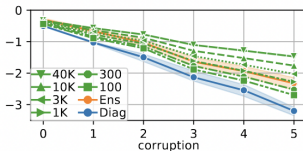
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(a) Rotated MNIST



(b) Corrupted CIFAR10

Subnet Size	Memory
11.2M (100%)	500TB
40K (0.36%)	6.4GB
1K (0.01%)	4.0MB
100 (0.001%)	40KB

(c) Memory Footprints

Figure 1: Subnet inference with Laplace approx. on ResNet-18.

BNNs are hard – integration

- The integrals in eqs. (1) and (2) are intractable!
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is easily approximated using Monte Carlo:

$$p(\mathbf{y}^* | \mathbf{x}^*) \approx \frac{1}{N} \sum_{n=1}^N p(\mathbf{y}^* | \mathbf{x}^*, \boldsymbol{\theta}^{(n)}), \quad \boldsymbol{\theta}^{(n)} \sim p(\boldsymbol{\theta} | \mathcal{D}). \quad (3)$$

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- However, approximating the **posterior** distribution

$$p(\boldsymbol{\theta} | \mathcal{D}) = \frac{p(\mathcal{D} | \boldsymbol{\theta}) p(\boldsymbol{\theta})}{\int_{\boldsymbol{\theta}} p(\mathcal{D} | \boldsymbol{\theta}) p(\boldsymbol{\theta}) d\boldsymbol{\theta}} \quad (1)$$

is slightly trickier...

Approximating the posterior

Two main approaches:

- 1 Assuming a simplified form for the posterior $p(\theta | \mathcal{D})$, allowing us to avoid (or simplify) calculating the evidence $p(\mathcal{D})$.

Approach 1 – Simplified Posteriors

- Laplace approx. [MacKay, 1992, Daxberger et al., 2021a].
- VI [Hinton and Van Camp, 1993, Graves, 2011, Blundell et al., 2015, Osawa et al., 2019].
- EP [Hernández-Lobato and Adams, 2015].
- MC Dropout [Gal and Ghahramani, 2016].
 - Some issues [Osband, 2016].

Approximating the posterior

Two main approaches:

- 1 Assuming a simplified form for the posterior $p(\theta | \mathcal{D})$, allowing us to avoid (or simplify) calculating the evidence $p(\mathcal{D})$.
- 2 Using MCMC methods to sample directly from $p(\theta | \mathcal{D})$ without ever calculating $p(\mathcal{D})$.

Approach 2 – Sampling

- Pioneered by Neal [1995], who used HMC [Duane et al., 1987, Neal, 2012]. “Gold standard”.
- SGLD [Welling and Teh, 2011] & SGHMC [Chen et al., 2014].
 - Biased [Betancourt, 2015].
 - No rejection sampling [Garriga-Alonso and Fortuin, 2021].

A simple baseline – MFVI

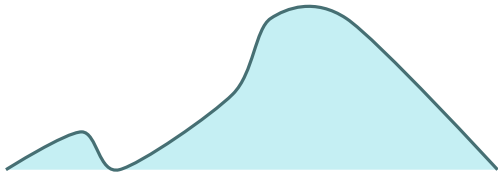
High level idea: approximate $p(\theta | \mathcal{D})$ with $q_\phi(\theta)$.

- q is a NN parameterised by ϕ .
- Mean-field assumption: dimensions of θ are independent.

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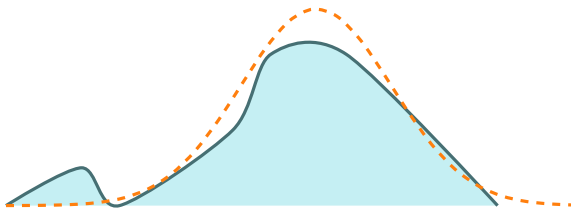
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$$\phi = \arg \min_{\phi} D_{\text{KL}} [q_\phi(\theta) || p(\theta | \mathcal{D})] \quad (4)$$

$$= \arg \max_{\phi} \mathbb{E}_{q_\phi(\theta)} [\log p(\mathcal{D} | \theta)] - D_{\text{KL}} [q_\phi(\theta) || p(\theta)] \quad (5)$$

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- 2 $D_{\text{KL}} [q_{\phi}(\boldsymbol{\theta}) || p(\boldsymbol{\theta})]$ – complexity term

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$$\mathcal{L}(\phi) \approx \frac{1}{N} \sum_{n=1}^N \left[\log p(\mathcal{D} | \boldsymbol{\theta}^{(n)}) - \log q_{\phi}(\boldsymbol{\theta}^{(n)}) + \log p(\boldsymbol{\theta}^{(n)}) \right],$$
$$\boldsymbol{\theta}^{(n)} \sim q_{\phi}(\boldsymbol{\theta}). \quad (7)$$

Problems with MFVI for BNNs

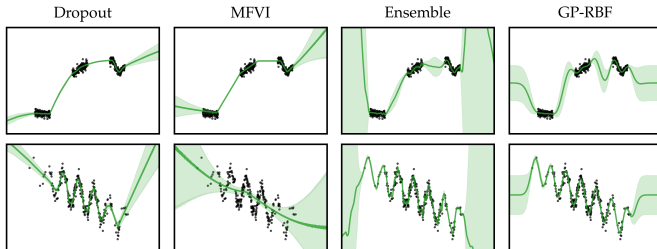


Figure 2: MFVI doesn't provide "in-between" uncertainty, and underfits!

Problems with MFVI for BNNs

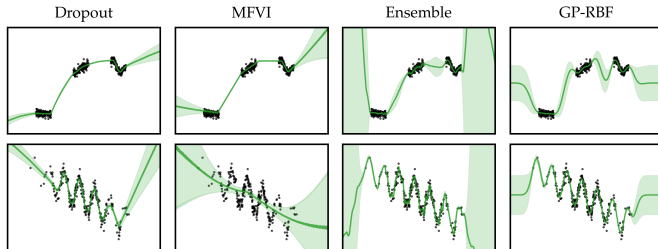


Figure 2: MFVI doesn't provide “in-between” uncertainty, and underfits!

- Foong et al. [2020] prove that MFVI (and MC Dropout) cannot capture “in-between” uncertainty for single hidden layer BNNs.
- They demonstrate this is a problem of approximate inference.
- They show empirically that this also occurs for deeper BNNs (despite proving that they are universal approximators for μ and σ).
- Farquhar et al. [2020] argue that MFVI is less restrictive with depth.

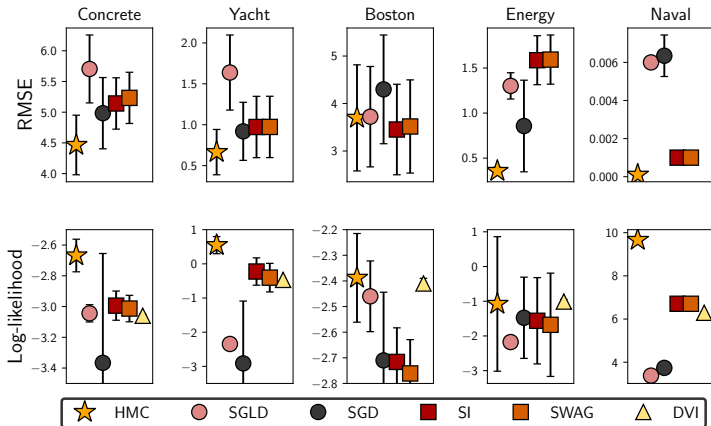
What are BNN posteriors really like?

Izmailov et al. [2021b] perform *full batch* 🔥 HMC for modern NNs to explore this question.

Note: this is not practical at all! But we can learn a lot.

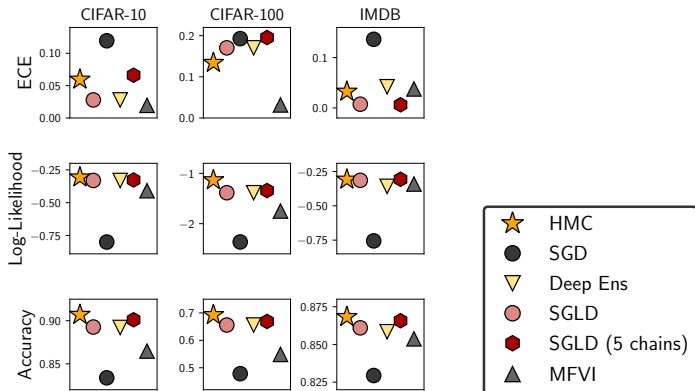
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Finding 1 – BNNs can achieve significant performance gains



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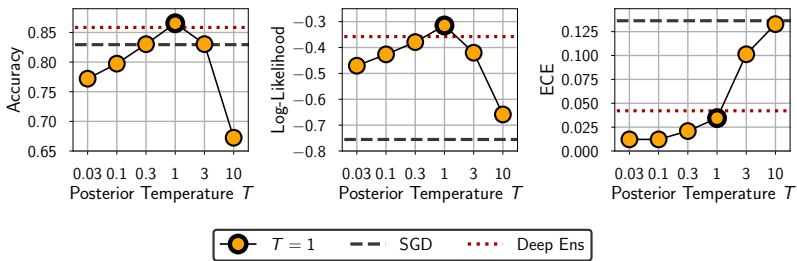


What are BNN posteriors really like?

Finding 2 – Posterior tempering is not needed

There is little evidence for a “cold posterior” effect [Wenzel et al., 2020], which seems to be largely caused by data augmentation.

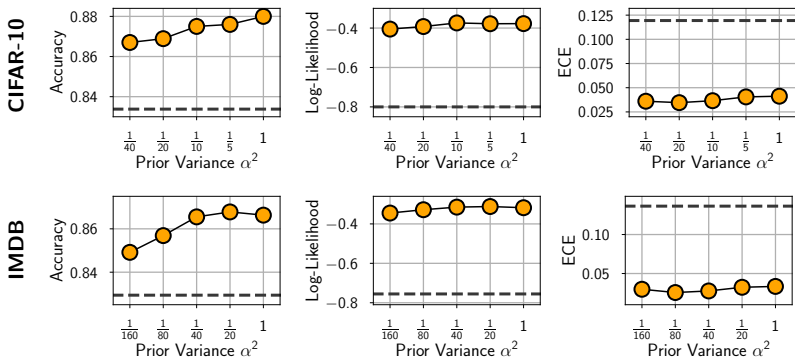
$$p_T(w|\mathcal{D}) \propto (p(\mathcal{D}|w) \cdot p(w))^{1/T} \quad (8)$$



What are BNN posteriors really like?

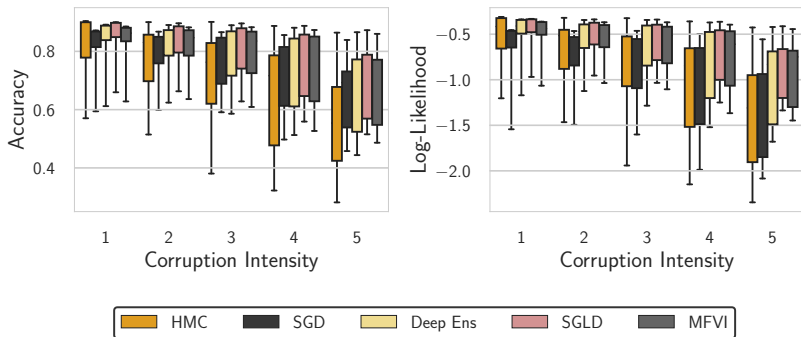
Finding 3 – Performance is robust to the choice of prior scale

... and pretty similar for diag. Gaussian, MoG, and logistic priors.



What are BNN posteriors really like?

Finding 4 – BNNs are surprisingly bad under dist. shift



Addressed in [Izmailov et al., 2021a] with prior choice.

What are BNN posteriors really like?

Finding 5 – Deep ens. and SGMCMC provide distinct predictive dists. from HMC

Deep ensembles and SGMCMC can provide good generalization. Deep ensemble predictive distributions are as to HMC as SGLD, and closer than VI.

Metric	HMC (reference)	SGD	Deep Ens	MFVI	SGLD	SGHMC	SGHMC CLR	SGHMC CLR-Prec
CIFAR-10								
Accuracy	89.64 ± 0.25	83.44 ± 1.14	88.49 ± 0.10	86.45 ± 0.27	89.32 ± 0.23	89.38 ± 0.32	89.63 ± 0.37	87.46 ± 0.21
Agreement	94.01 ± 0.25	85.48 ± 1.00	91.52 ± 0.06	88.75 ± 0.24	91.54 ± 0.15	91.98 ± 0.35	92.67 ± 0.52	90.96 ± 0.24
Total Var	0.074 ± 0.003	0.190 ± 0.005	0.115 ± 0.000	0.136 ± 0.000	0.110 ± 0.001	0.109 ± 0.001	0.099 ± 0.006	0.111 ± 0.002

What are BNN posteriors really like?

(In an idealised setting...)

- 1 BNNs can achieve significant performance gains over standard training and deep ensembles.
- 2 Posterior tempering is not needed for near-optimal performance, with little evidence for a “cold posterior” effect (largely caused by data augmentation).
- 3 Performance is robust to the choice of prior scale, and relatively similar for diagonal Gaussian, MoG, and logistic priors.
- 4 BNNs show surprisingly poor generalization under distribution shift. Addressed in [Izmailov et al., 2021a] with prior choice.
- 5 Deep ensembles and SGMCMC can provide good generalization, but different predictive distributions from HMC. Notably, deep ensemble predictive distributions are as to HMC as SGLD, and closer than VI.

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